

Initiation and early development of the 2008 April 26 CME

J. Huang

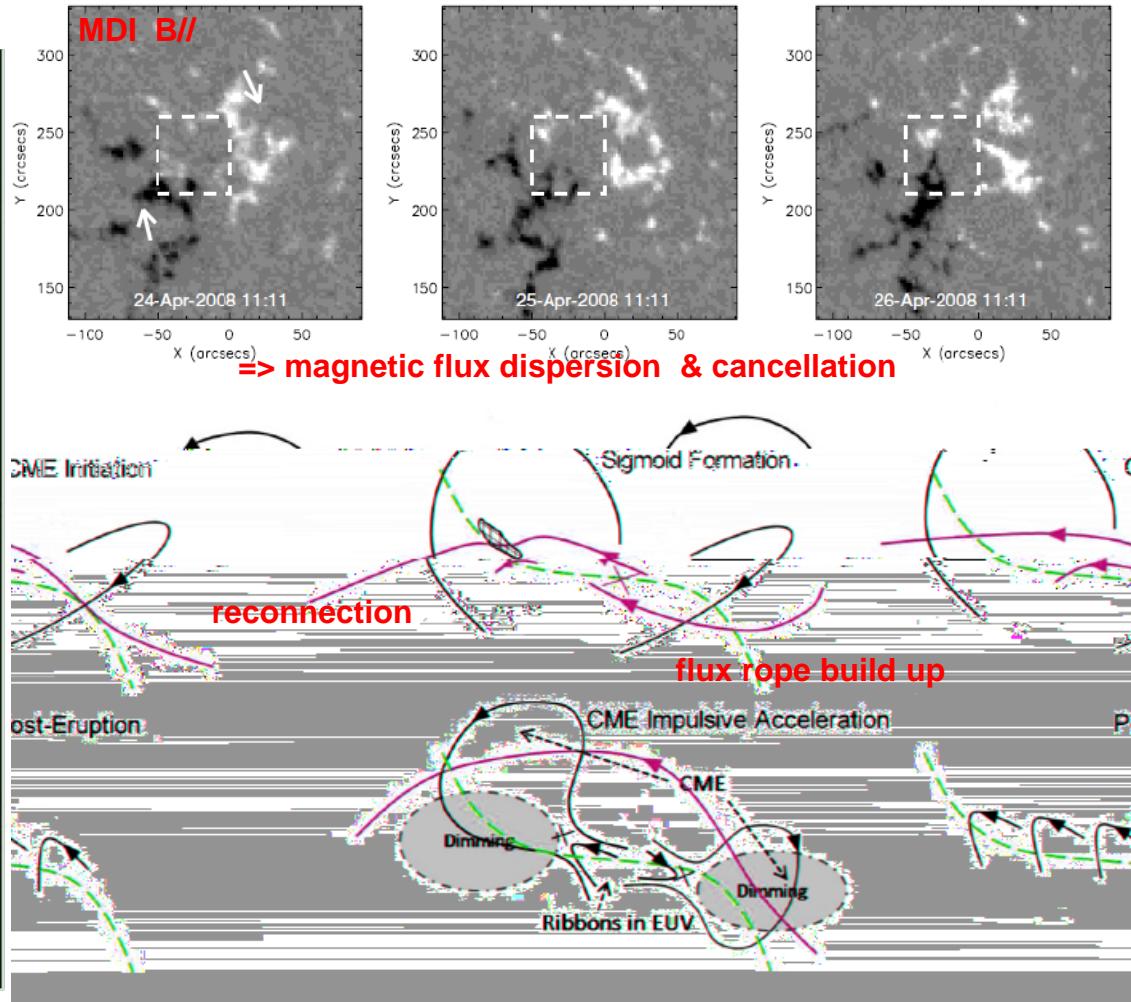
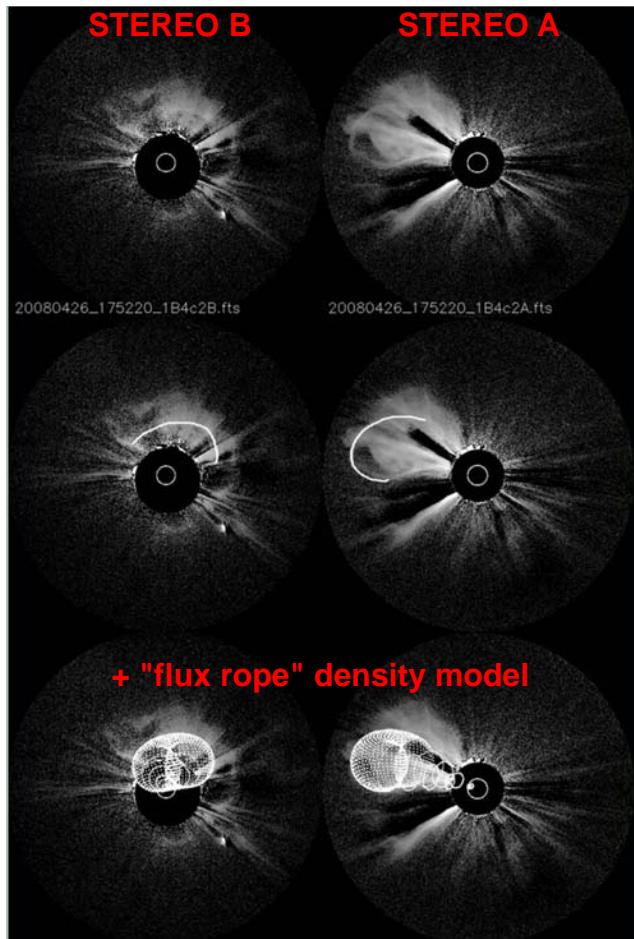
National Astronomical Observatories, CAS, Beijing, China

P. Démoulin, M. Pick

LESIA, Observatoire de Paris, Meudon, France

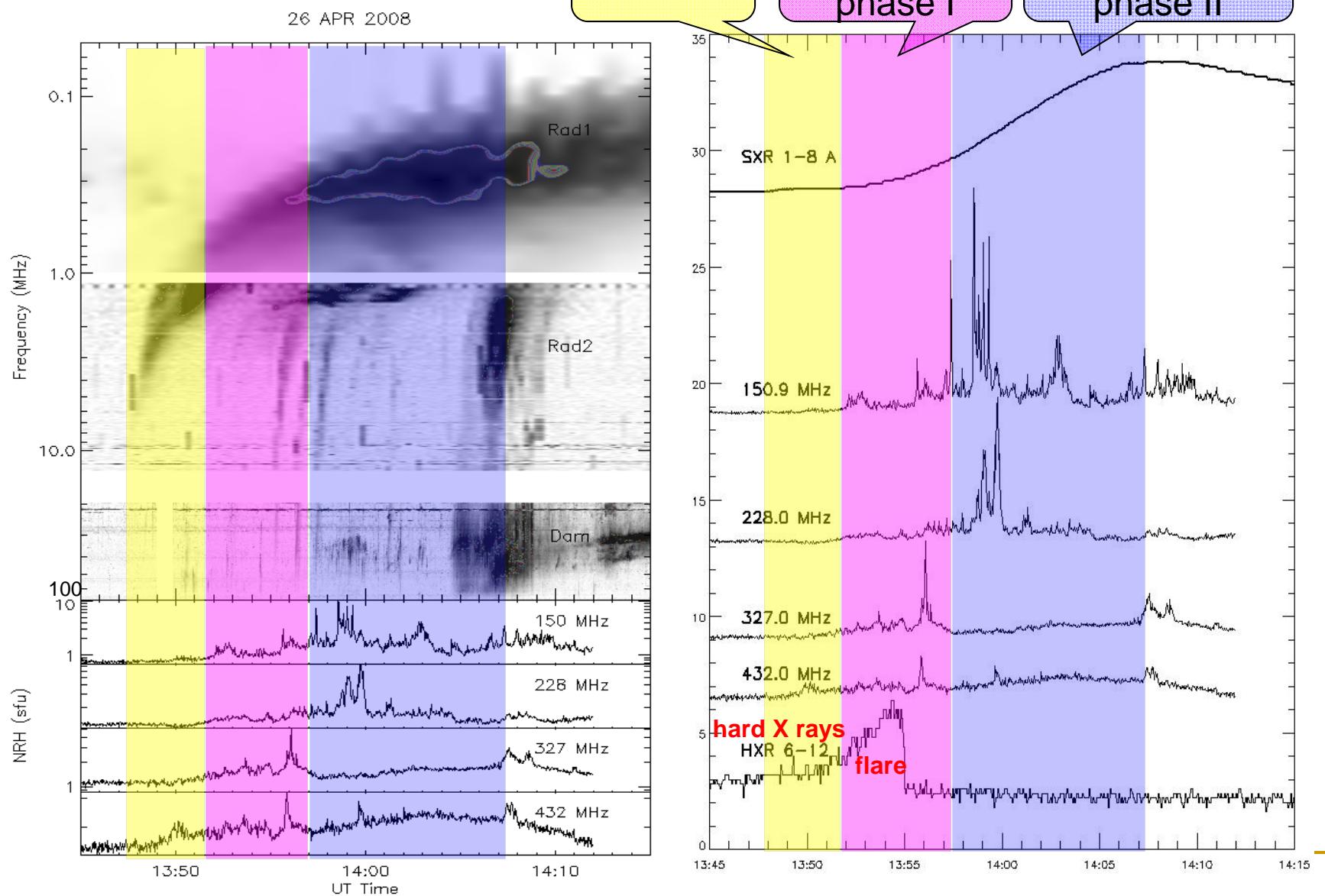
Introduction:

2008 April 26 event



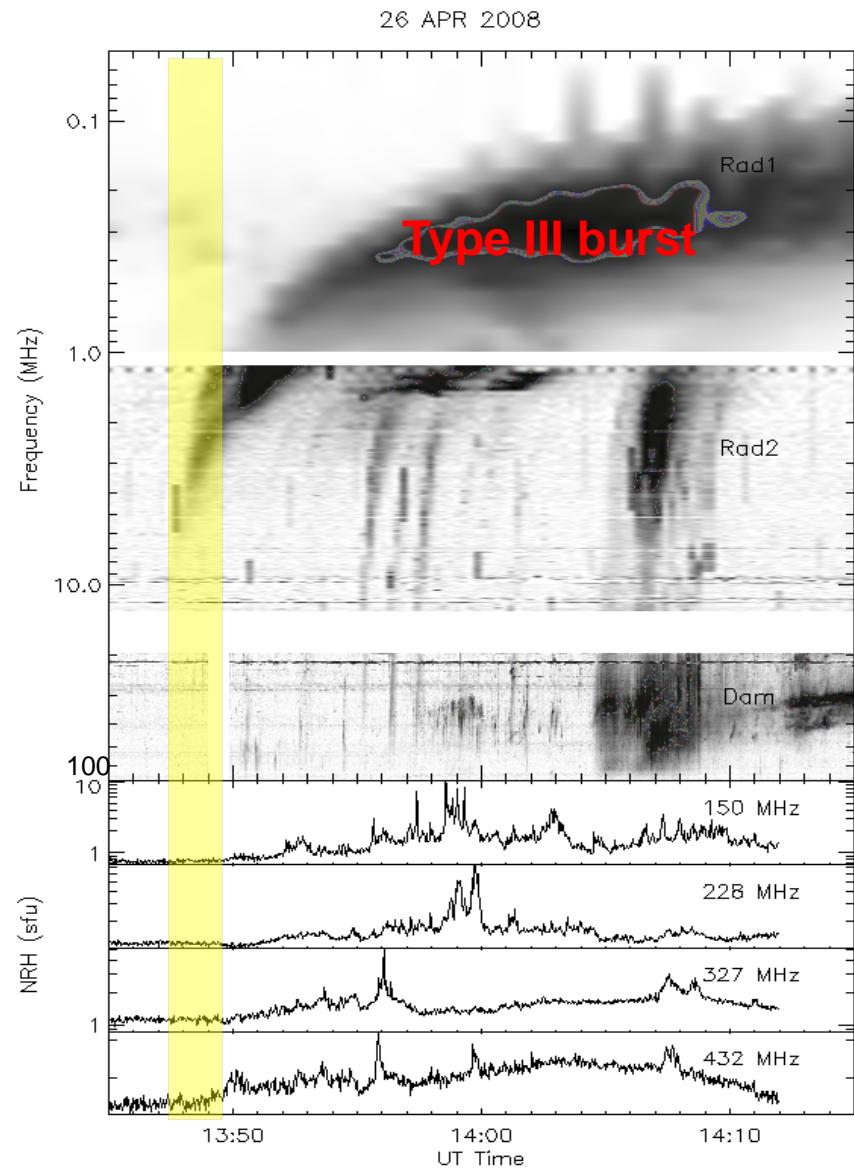
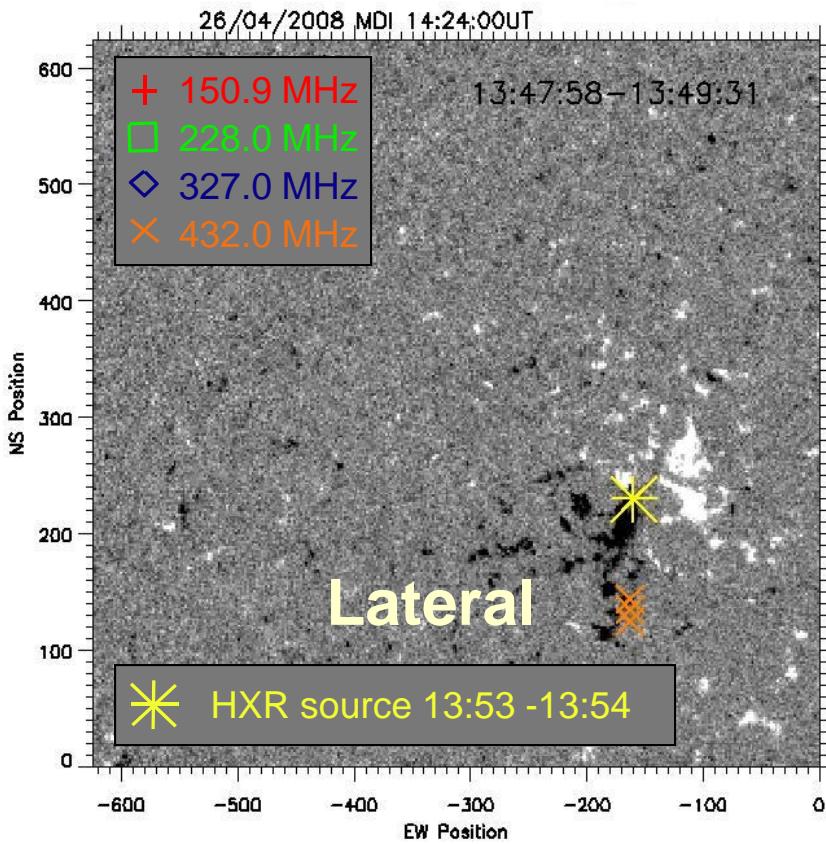
(Thernisien et al. 2009, Wood et al. 2009)

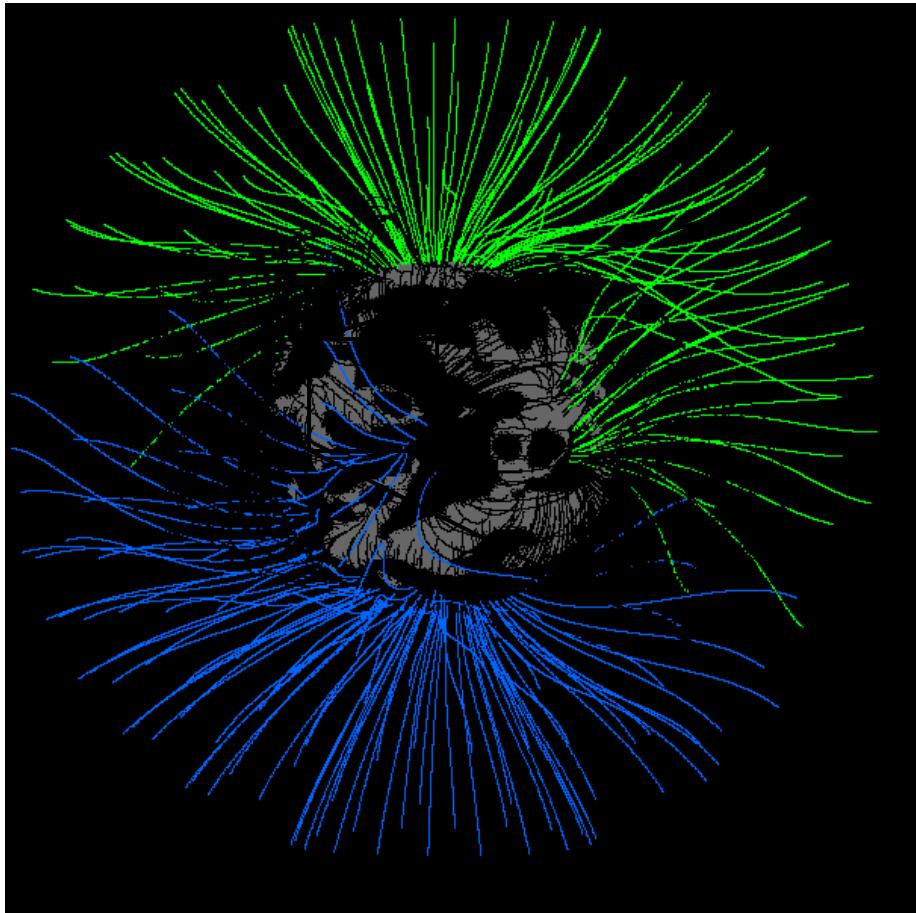
The radio burst:



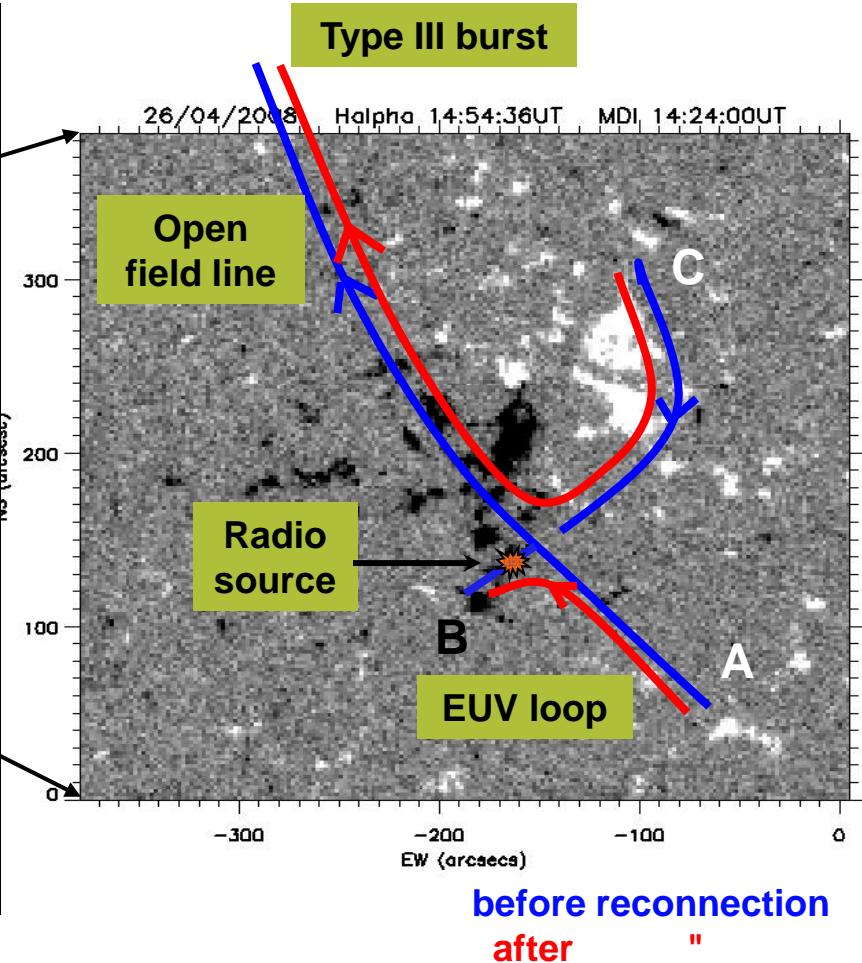
Initiation I:

■ 13:47:58-13:49:31



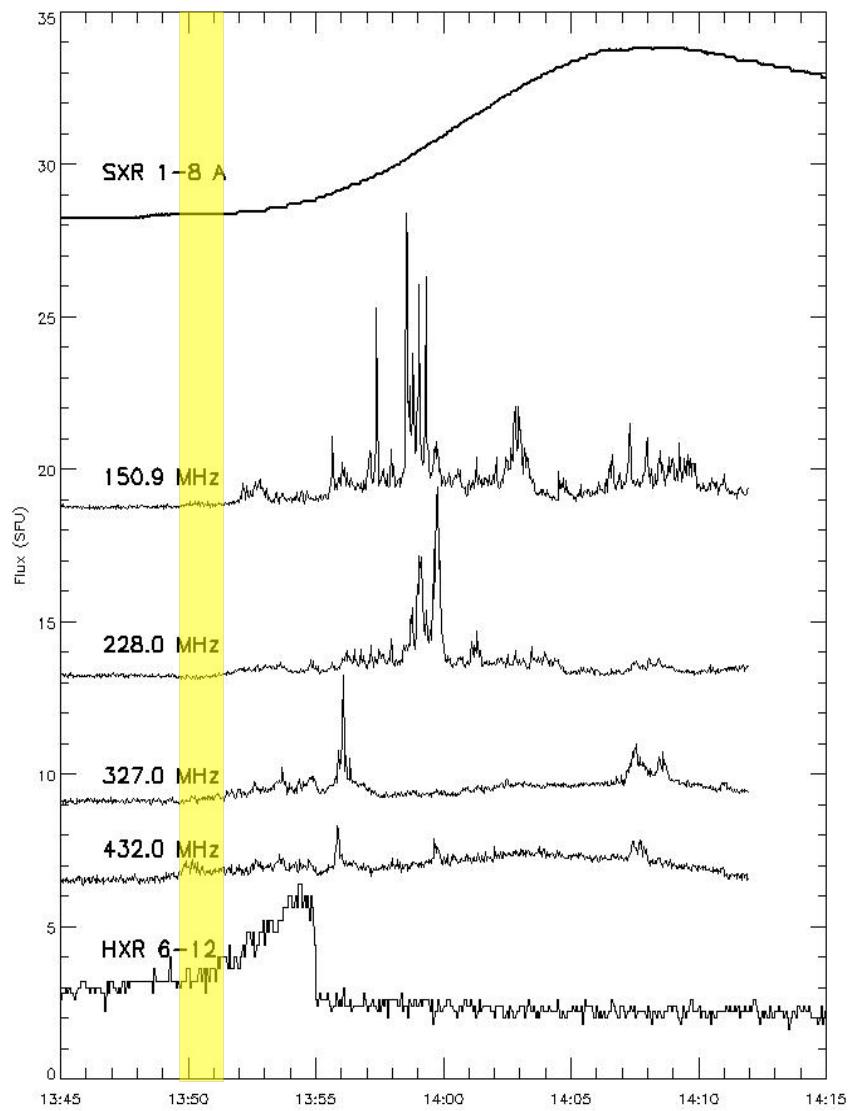
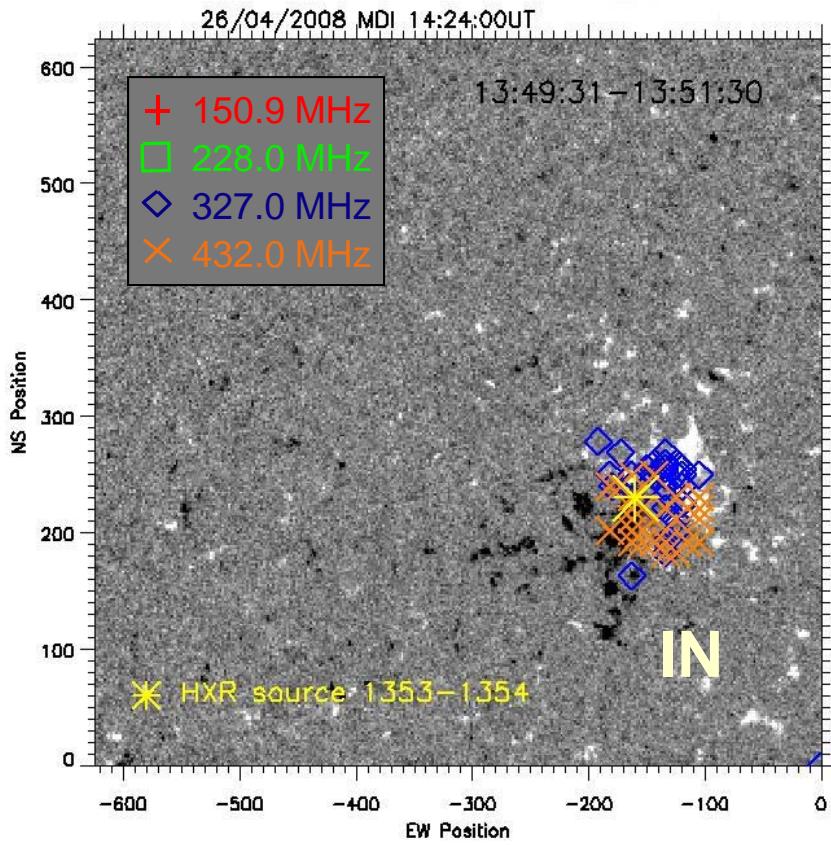


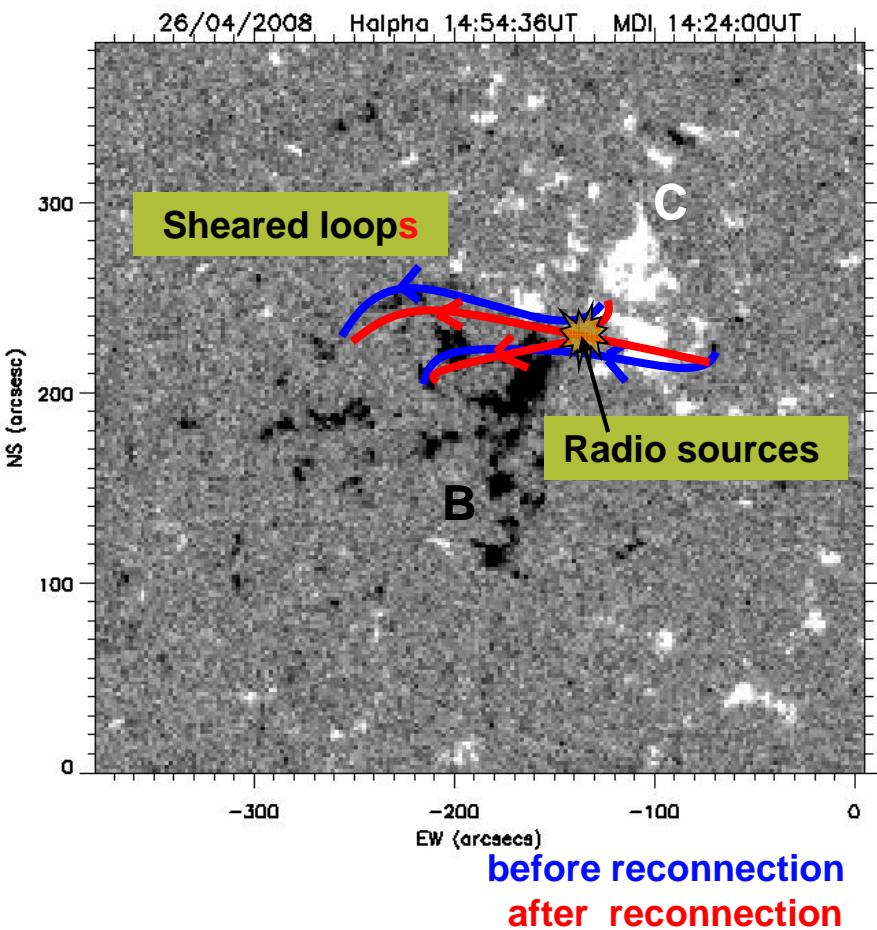
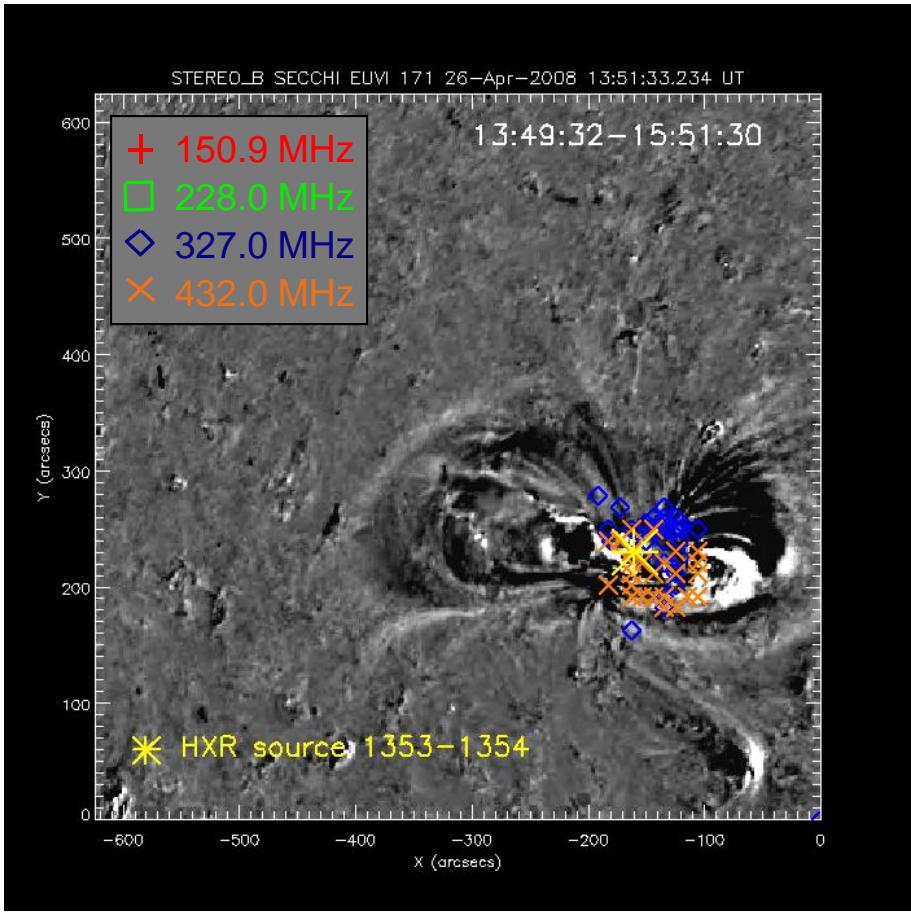
Yiming Wang



Initiation II:

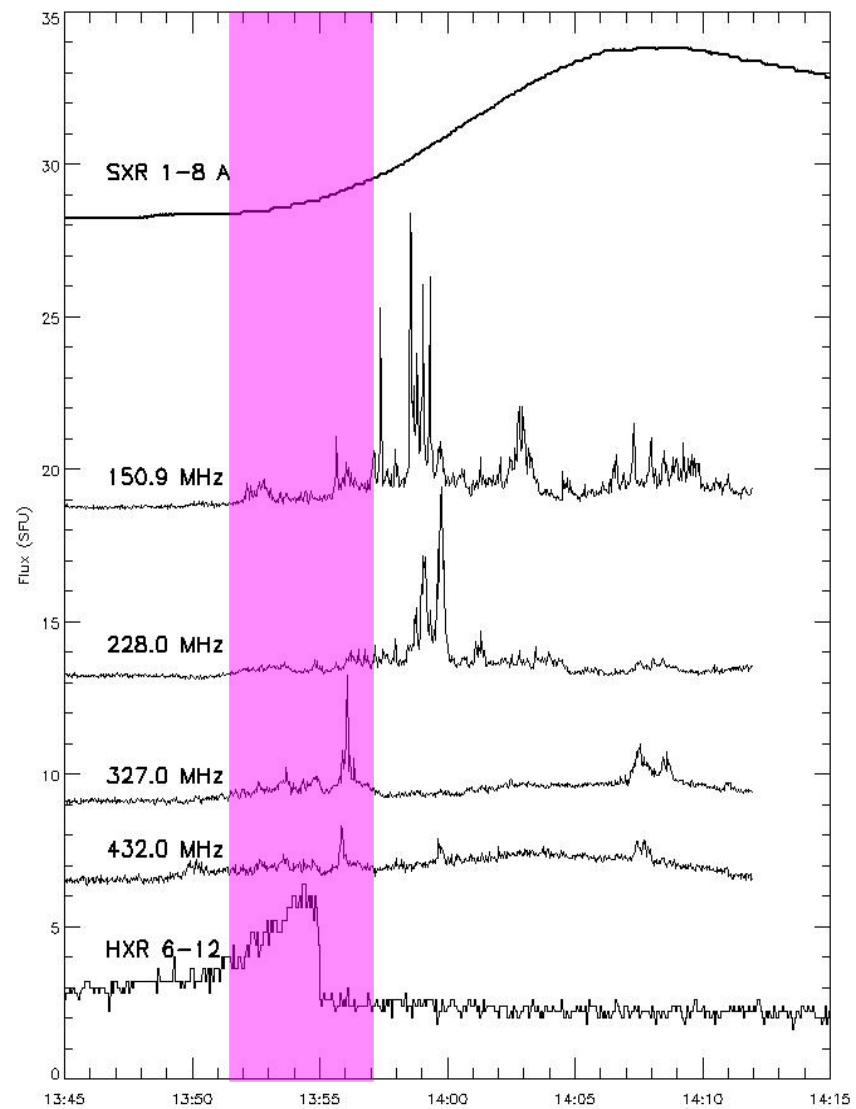
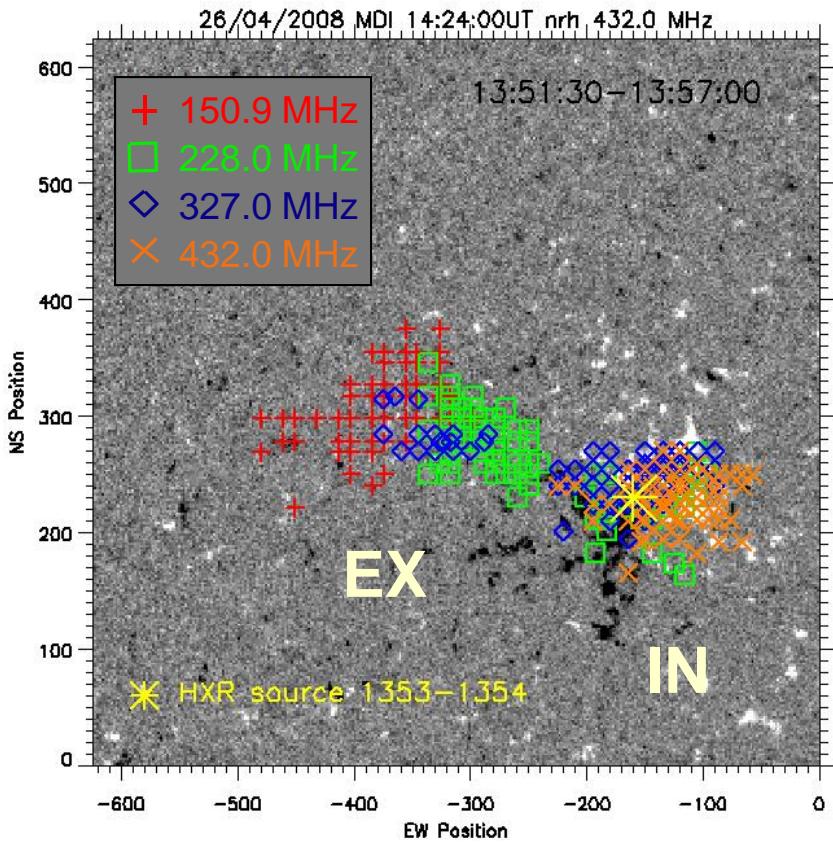
■ 13:49:31-13:51:30



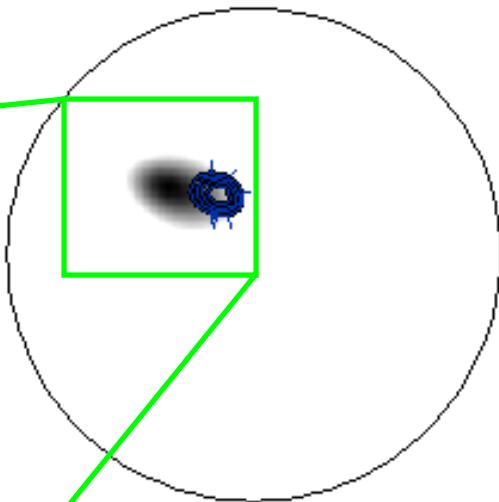
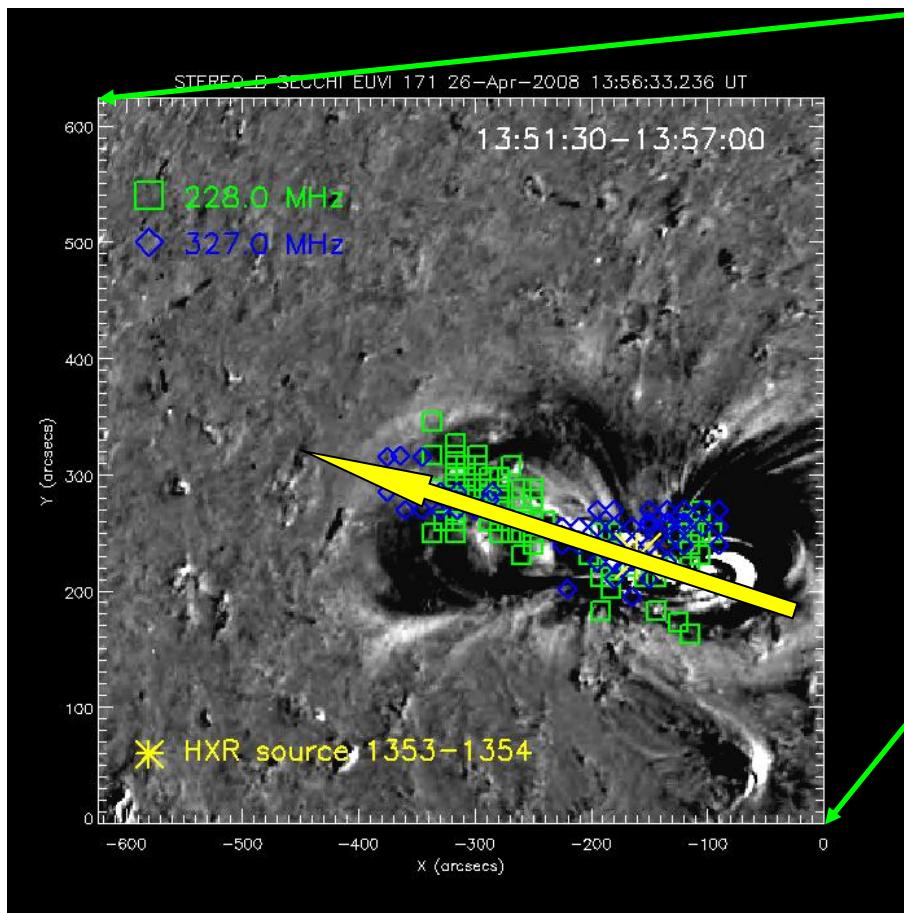


Eruptive phase I:

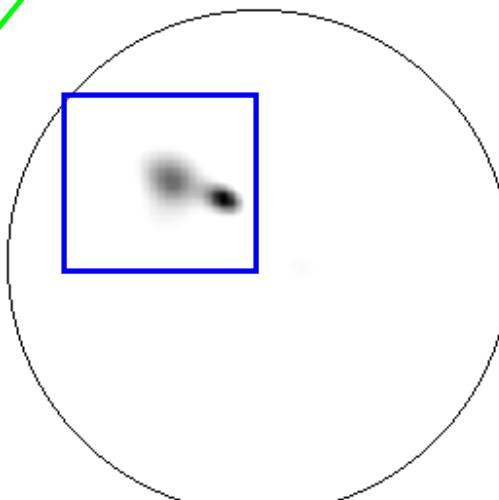
■ 13:51:30-13:57:00



228 MHz 13:55:36 13:55:42



Nancay

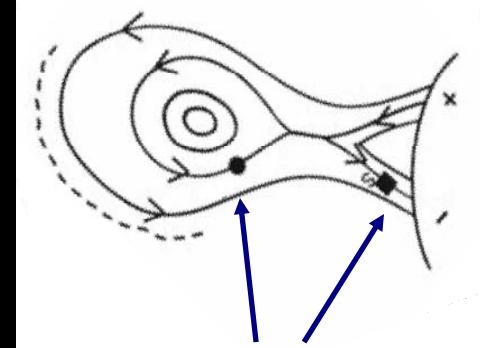
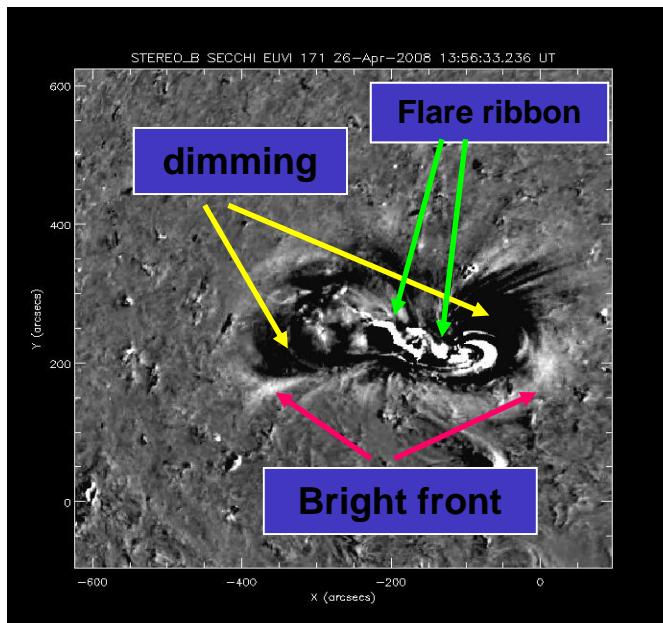
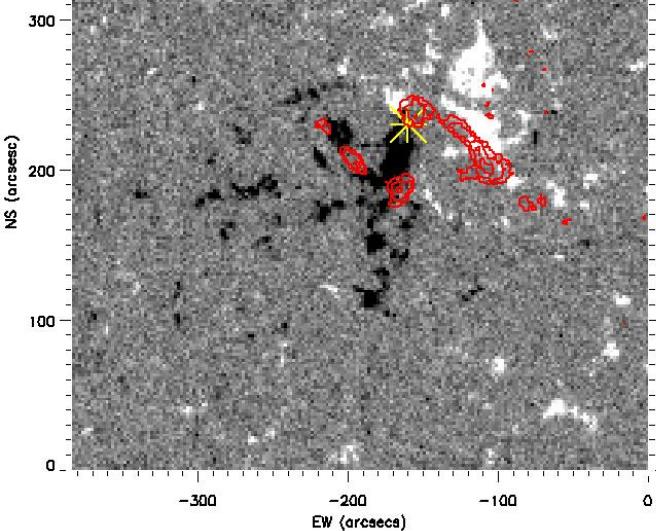


The double sources and dimmings indicate the direction of the flux rope.

327 MHz 13:52:26

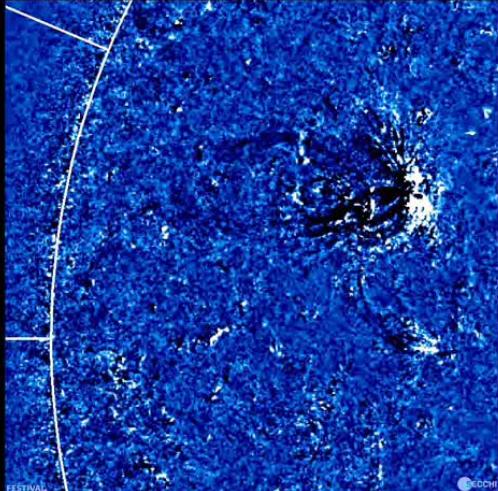
26/04/2008 Halpha 14:54:36UT MDI 14:24:00UT

* HXR source 1353-1354
Red --- Halpha



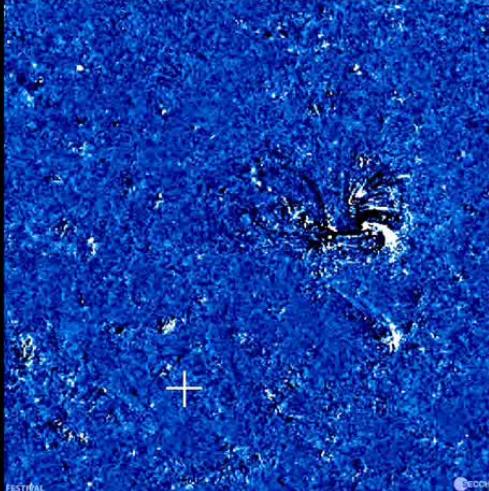
Double radio sources

A EUVI:08-04-26 13:33:30



Stereo A

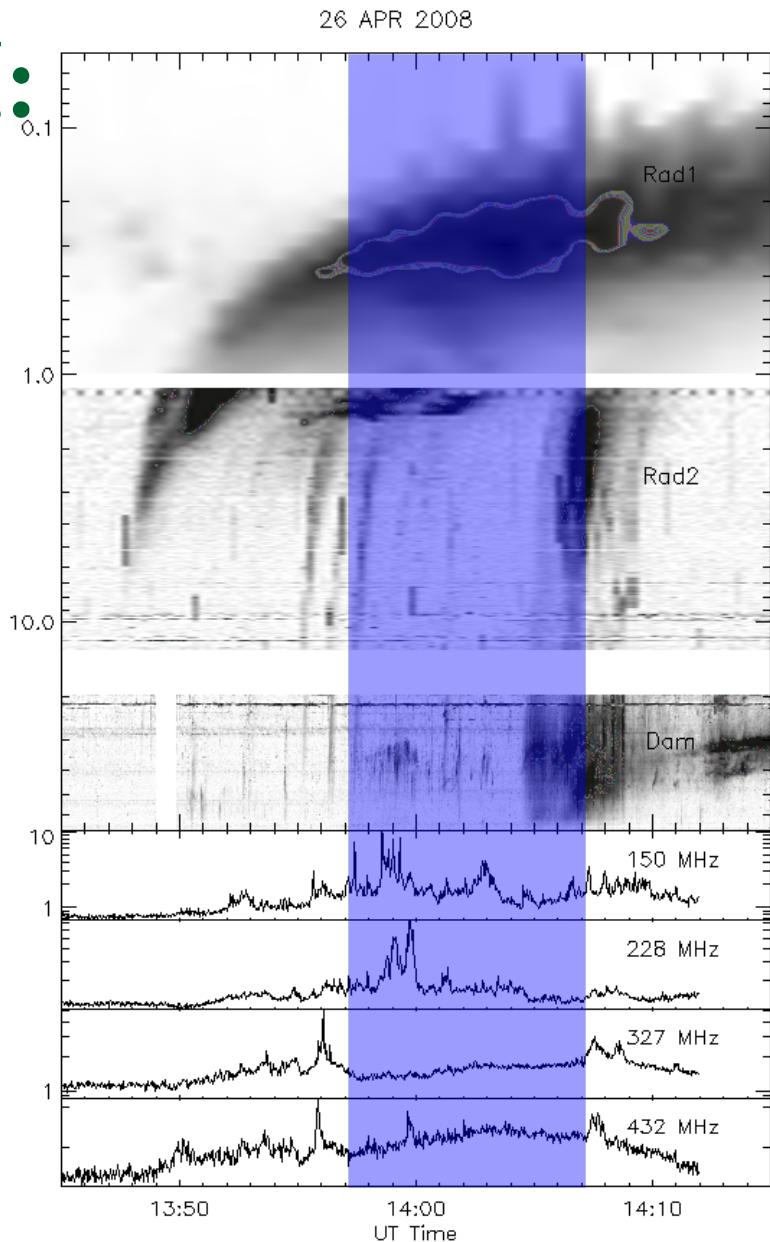
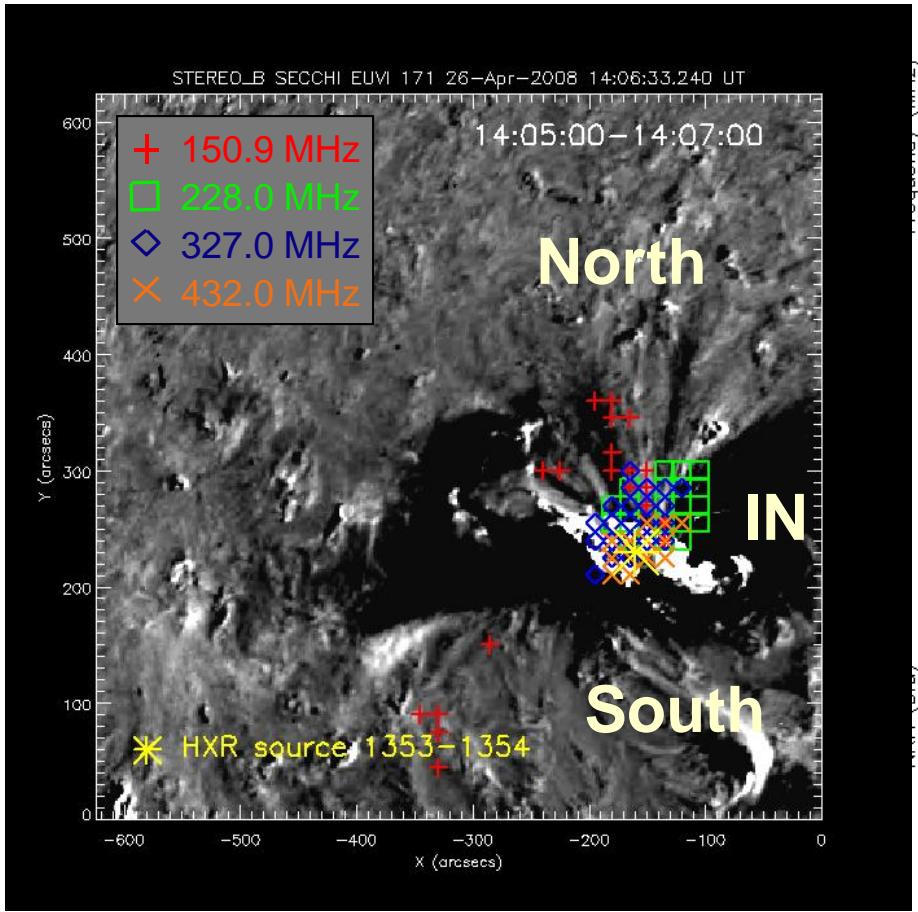
B EUVI:08-04-26 13:34:03



Stereo B

Eruptive phase II:

■ 13:57:00-14:07:00



Summary:

- Initiation I (13:47-13:50)

The lateral source: traces reconnection with open field line => type III burst
=> reduction of the downward magnetic tension

- Initiation II (13:50-13:51)

The IN source: new reconnection of sheared loops occurred over the PIL
=> build up of the flux rope

- Eruptive phase I (13:51-13:57)

The IN source (over PIL): energetic electrons in the flare-loop top
The EX source (over dimming): energetic electrons at the flux-rope bottom
gives the direction of the eruptive flux rope

- Eruptive phase II (13:57-14:07)

The south and north sources: preceding the EUV bright front
trace the expansion of the flux rope

