

# Forecasting the magnetic disturbance-storm-time ( $Dst$ ) index using machine-learning

Manoj Nair, Patrick Alken and Arnaud Chulliat

# Speaker Bio

- Manoj Nair
  - Research scientist and Operational science-lead
  - University of Colorado and US National Oceanic and Atmospheric Administration
  - 17+ years research experience in geomagnetism
  - PhD in Geophysics
  - Boulder, CO
  - Specialized in
    - Geomagnetism
      - Signal-processing, research-to-operations
      - Machine-learning

# Geomagnetism group of University of Colorado and US National Oceanic and Atmospheric Administration

- Conducts original research on geomagnetism
- Develops and distributes magnetic reference models (HDGM, WMM, IGRF)
- Real-time modeling of magnetic disturbance field
- Magnetic survey data repository (GEODAS)
- CrowdMag citizen-science project

<https://geomag.colorado.edu/>

<https://www.ngdc.noaa.gov/geomag/>



# Roadmap

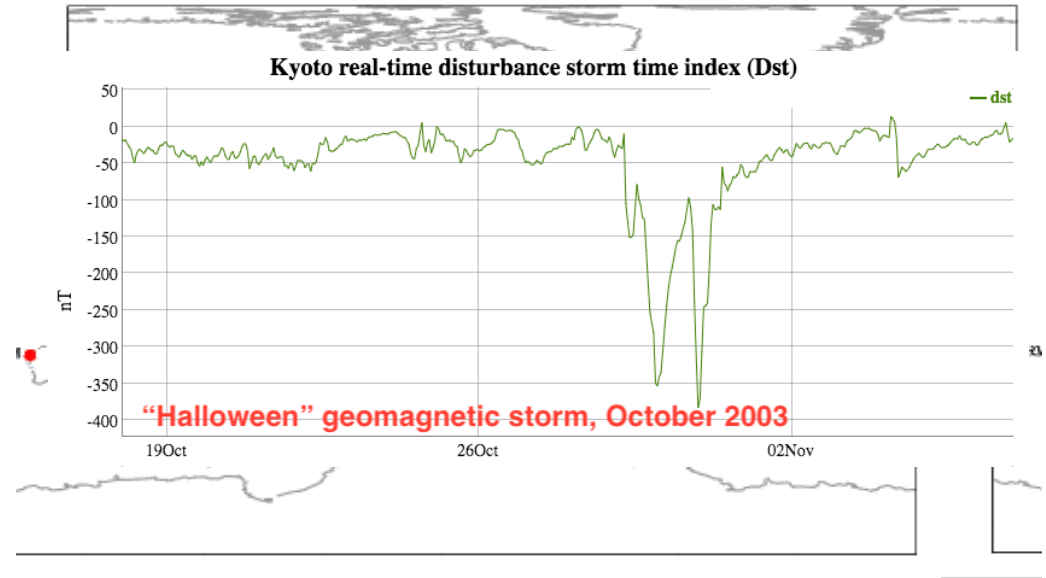
- Disturbance-storm-time ( $Dst$ ) index: what and why
- Solar-wind based forecast of  $Dst$
- Machine-Learning and artificial neural networks
- Modeling of  $Dst$
- Results and conclusion

4

# Disturbance-storm-time (*Dst*) index

A measure of magnetic disturbance

- Solar-wind interaction with Earth's magnetic field generate electric currents
- *Dst* index is a measure of "ring-currents" in the magnetosphere
- Hourly *Dst* index is calculated using four geomagnetic observatories
- Different flavors: Kyoto *Dst*, USGS *Dst*, *Rc* index

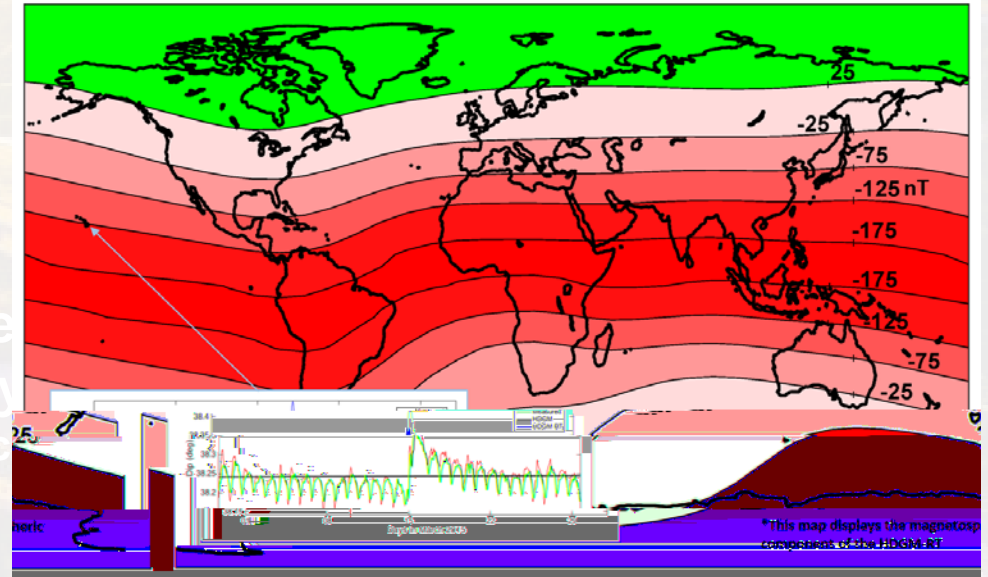


# Why to predict *Dst*?

## Important space-weather specification

6

- Ring-current is one of the major current systems in the magnetosphere
- Critical input to magnetospheric specification models
- Operational *Dst* forecast provides early warning
- Augment NOAA/CIRES real-time magnetic disturbance modeling



# Forecasting of $Dst$ using solar-wind data

- Solar-wind forecasting
  - Less-accurate
  - Lead-time
  - Observatory data not needed
- Empirical relationship
  - Burton et al (1975), Temerin and Li (2002), O'Brien and McPherron (2000)
- Physics-based models
  - University of Michigan's Geospace model
- Machine-learning approach

The  $dst1$ ,  $dst2$ , and  $dst3$  terms are calculated as follows:

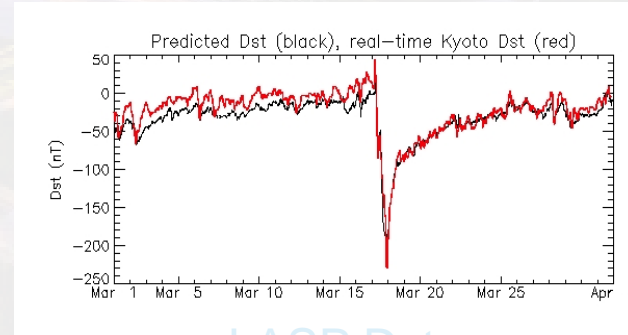
$$dst1(t+dt) = dst1(t) + \{a_1 \cdot [-dst1(t)]^{a_2} + fe1(t) [1 + \frac{a_3 \cdot dst1(t - \tau_1) + a_4 \cdot dst2(t - \tau_2)}{1 - a_5 \cdot dst1(t - \tau_1) - a_6 \cdot dst2(t - \tau_2)}]\} dt \quad (5)$$

$$dst2(t+dt) = dst2(t) + \{b_1 \cdot [-dst2(t)]^{b_2} + fe2(t) [1 + \frac{b_3 \cdot dst1(t - \tau_2)}{1 - b_5 \cdot dst1(t - \tau_2)}]\} dt \quad (6)$$

$$dst3(t+dt) = dst3(t) + \{c_1 \cdot dst3(t) + fe3(t) [1 + \frac{c_2 \cdot dst3(t - \tau_3)}{1 - c_2 \cdot dst3(t - \tau_3)}]\} dt \quad (7)$$

where  $a_1 = 6.51 \cdot 10^{-2}$ ,  $a_2 = 1.370$ ,  $a_3 = 8.4 \cdot 10^{-3}$ ,  $a_4 = 6.053 \cdot 10^{-3}$ ,  $a_5 = 1.21 \cdot 10^{-3}$ ,  $a_6 = 1.55 \cdot 10^{-3}$ ,  $\tau_1 = 0.14$  days,  $b_1 = 0.792$ ,  $b_2 = 1.326$ ,  $b_3 = 1.29 \cdot 10^{-2}$ ,  $b_5 = 0.18$  days,  $c_1 = -24.3$ ,  $c_2 = 5.2 \cdot 10^{-2}$ ,  $\tau_3 = 9 \cdot 10^{-2}$  days,  $fe1 = -4.96 \cdot 10^{-3} (1 + 0.28 \cdot dh) [2 \cdot \text{exx} + \text{abs}(\text{exx} - th1) + \text{abs}(\text{exx} - th2) - th1 - th2] v_x^{1.11} a^{0.89} \sin^{0.6}(\phi)$ ,  $fe2 = 2.02 \cdot 10^3 \cdot \sin^{13}(\phi) \cdot df2 / (1 - df2)$ ,  $df2 = -3.85 \cdot 10^{-8} \cdot v_x^{1.97} b_x^{1.16} \sin^{3.7}(\theta) \cdot a^{0.41} \cdot (1 + dh)$ ,  $fe3 = 3.45 \cdot 10^3 \cdot \sin^{0.9}(\theta) \cdot df3 / (1 - df3)$ ,  $df3 = -4.75 \cdot 10^{-6} \cdot v_x^{1.22} \cdot b_x^{1.11} \sin^{5.2}(\theta) a^{0.24} (1 + dh)$ ,  $\text{exx} = 10^{-3} \cdot v_x \cdot b_y \sin^{0.3}(\theta)$ ,  $\theta = -(\text{acos}(-\frac{b_x}{b_y}) - \pi) / 2$ ,  $b_x = (b_x^2 + b_y^2)^{1/2}$ ,  $th1 = 0.725 \sin^{-1.66}(\phi)$ ,  $th2 = 1.83 \sin^{-1.66}(\phi)$ ,  $dh = b_p \cdot \cos(\text{atan}(b_x, b_y) + 6.10) (3.59 \cdot 10^{-2} \cos(2\pi \cdot \text{yr} + 0.04) - 2.18 \cdot 10^{-2} \sin(2\pi \cdot \text{yr} - 1.60))$ , and  $b_p = (b_x^2 + b_y^2)^{1/2}$ . Time ( $t, dt$ ) is in days, magnetic field in  $nT$ , solar wind velocity in  $\text{km/s}$  and density in  $\text{cm}^{-3}$ . Here  $v_x$  is the magnitude of x-component of the solar wind velocity.

Temerin and Li (2002) – LASP model

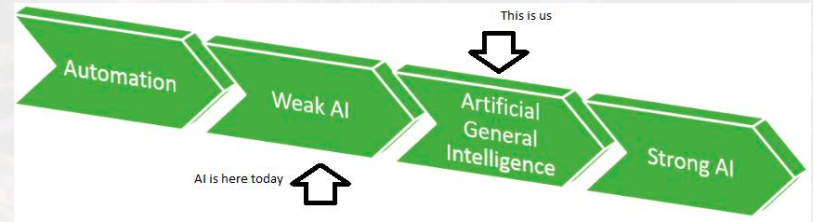


# Artificial Intelligence

- An “AI”, or Artificial Intelligence is an intelligent code/machine made by human.
- AI performs cognitive functions such as learning, problem solving, Planning.
- AI progression
  - Artificial Weak Intelligence
  - Artificial General Intelligence
  - Strong AI
- Practical applications are limited to Weak-AI
  - Machine-Learning



8



Source: <https://vincentlauzon.com/2015/09/16/strong-ai-existential-risks/>



# Machine-Learning

## Deep Neural Networks

### Artificial Neural Networks

- Mimics the function of brain
- Weights and transfer function
- Universal non-linear approximator
- Back-propagation training
- Supervised learning

### ML frameworks

- Bring your own software
- Tensorflow (Google)
- PyTorch (Facebook)

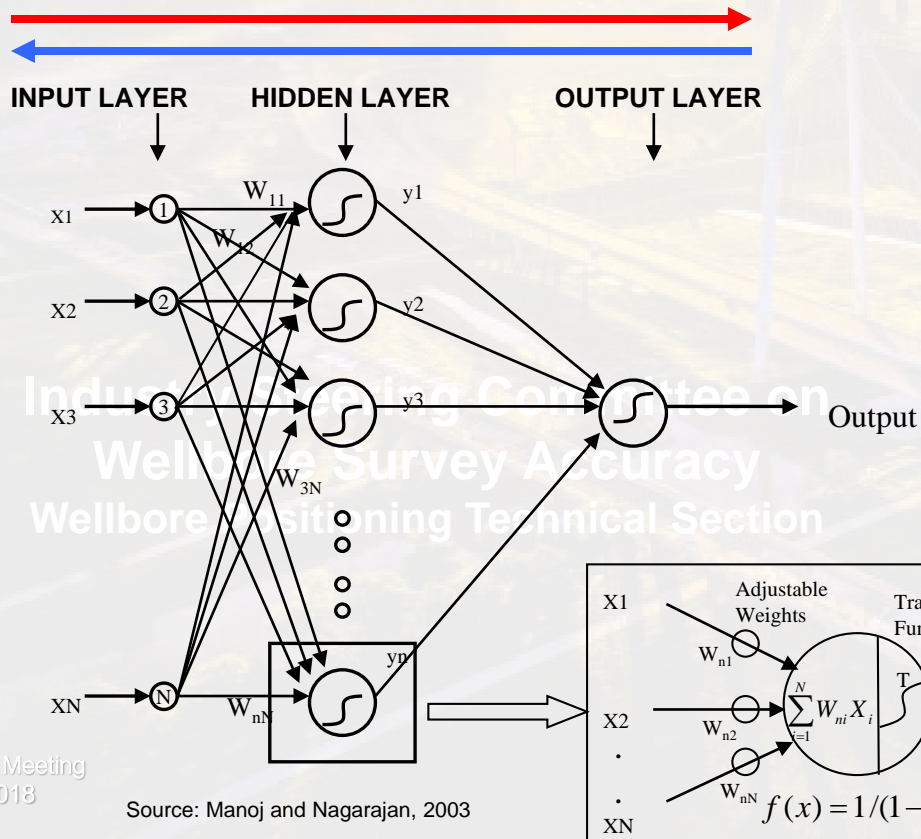
### Recurrent Neural Networks

9

- A variation of ANN
- For predicting temporal (sequential) information
- Sequence-to-sequence processing



# Training Deep Neural Networks



Source: Manoj and Nagarajan, 2003

Feed forward

$$y_j = f\left(\sum W_{ji}x_i\right)$$

$$f(x) = 1/(1 - e^{-x})$$

Error Backpropagation

$$\delta_j = \partial E / \partial W_{ij}$$

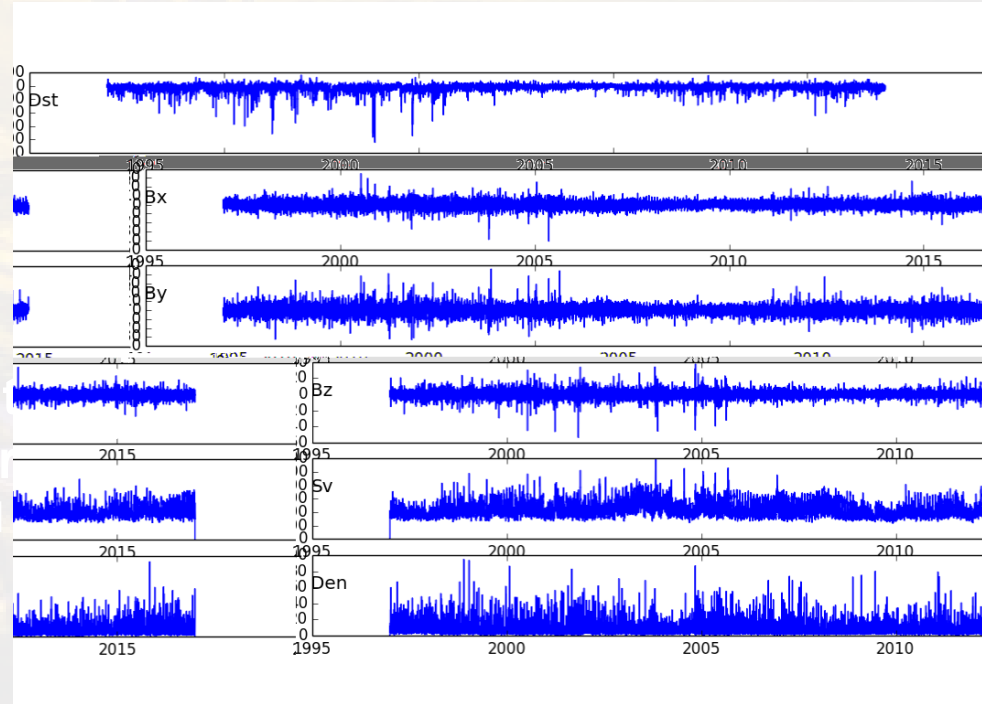
$$E = 0.5 \sum (d_j - o_j)^2$$

$$\Delta W_{ji} = \eta \delta_j o_i$$

$$\Delta W_{ji}(n+1) = \eta \delta_j o_i + \beta \Delta W_{ji}(n)$$

# Data used for machine-learning

- Observed *Dst* values (Kyoto WDC)
- Observed solar-wind data (NASA-OMNI)
- 1997-2016 (175,200 hourly values)
- Divided into training and testing segments
- Normalized



11

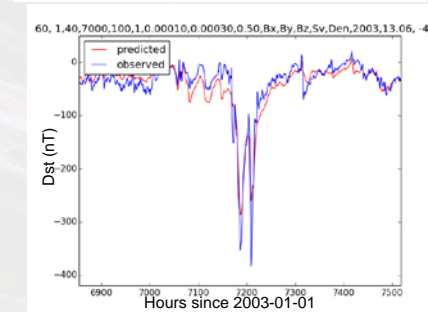
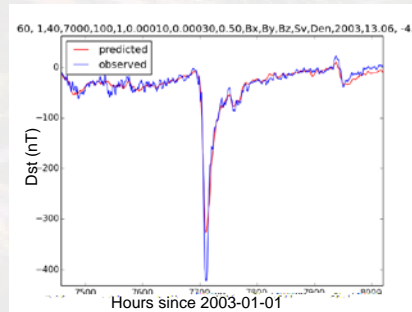
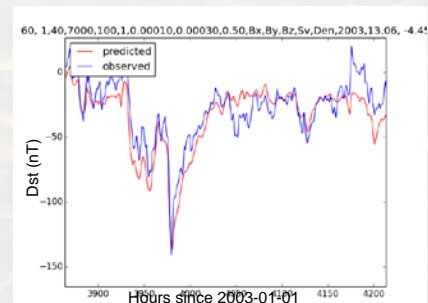
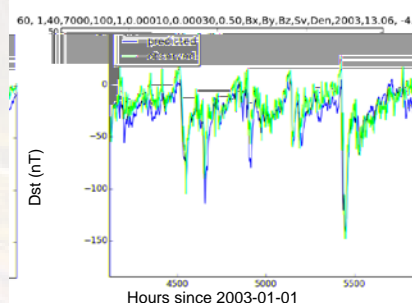
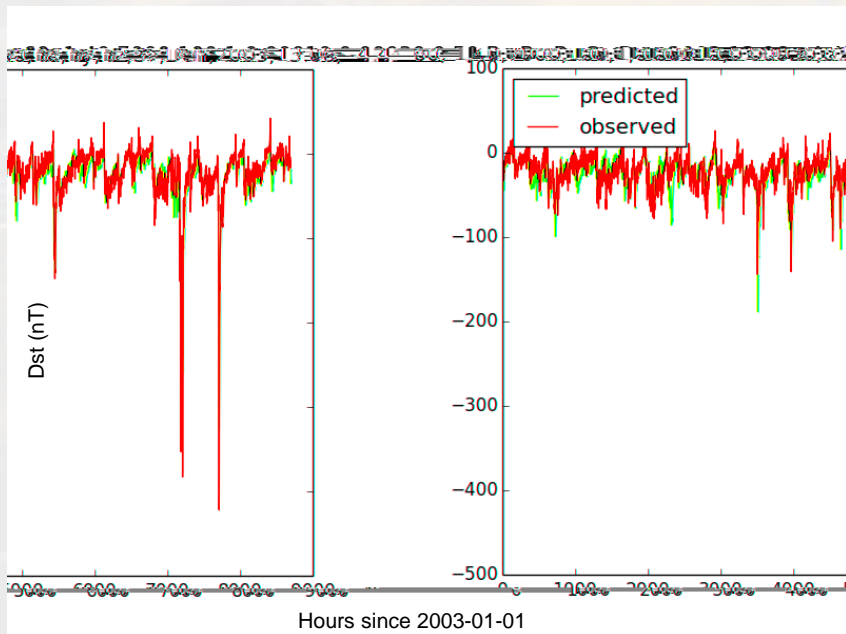
# Training the model

12

- Aim: one-step (hour) ahead forecast of  $Dst$  using current and historical solar-wind data
- Optimizing hyperparameters
- Minimizing the loss-function versus generalizing the model
- CPU versus GPU
- Final weights and biases saved for production.

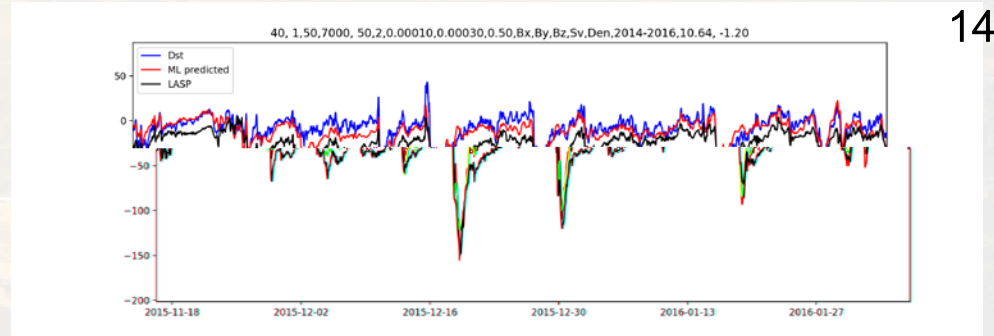
# Results

13

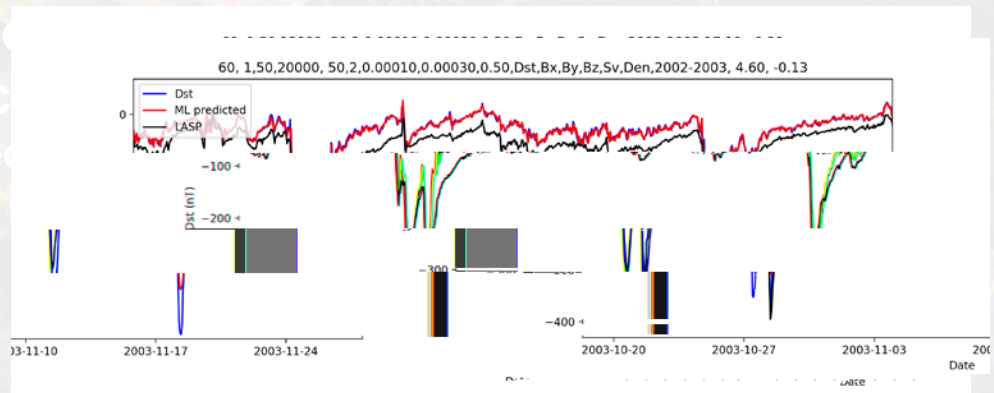


# Benchmarking ML prediction

- Compared ML model against LASP model using test data.
- ML and LASP predictions are very similar
- Extreme geomagnetic storm of Nov-2003 is better predicted by ML
- Further improvement to prediction is achieved by ingesting past *Dst* data

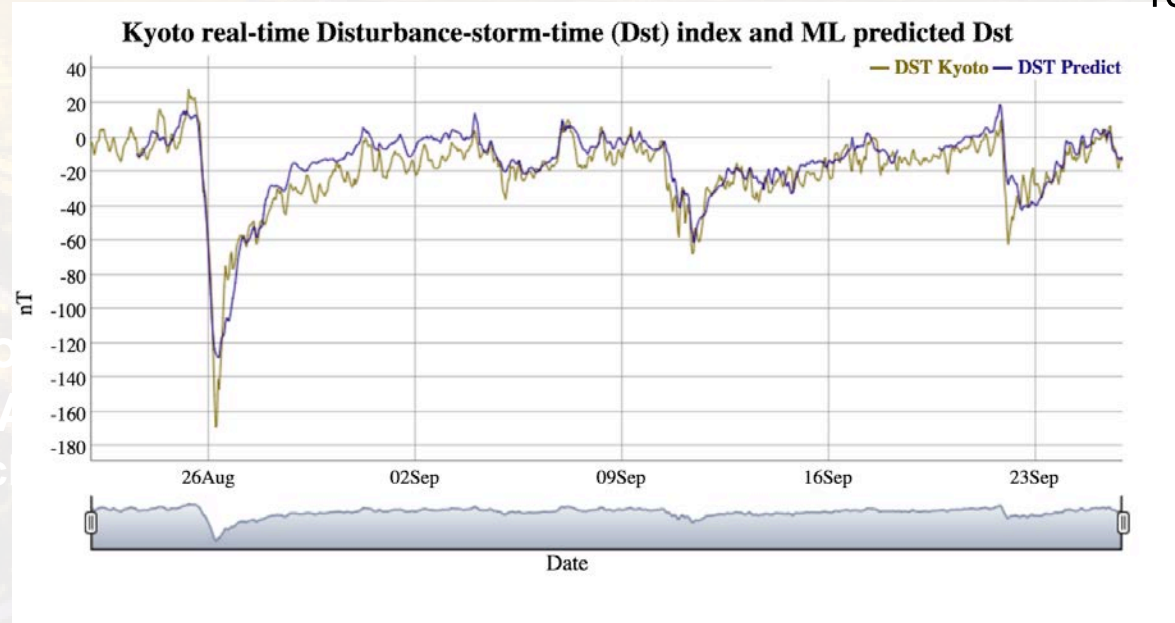


14



# Real-time prediction service

- We use the trained ML model to predict  $Dst$  in real-time
  - Satellite only
  - Satellite + past  $Dst$
- Uses NOAA's DSCOVR satellite data
  - Operational upstream solar-wind measurements
- 1-hour advance prediction of  $Dst$
- Real-time validation against observed data



Will be available at <https://geomag.colorado.edu>



# Conclusion

- Machine-learning (ML) is a powerful tool to develop predictive models
- Disturbance-storm-time (*Dst*) index is an important specification of magnetic disturbance
- Using historical *Dst* and satellite data, we developed a ML model to forecast *Dst* data
- Our predictions compare favorably with observed data
- Potential for modeling other electric-current systems in the space.